

Anisotropies in the gravitational wave background as a probe of the cosmic string network

based on arXiv:1604.00332

Sachiko Kuroyanagi

(Nagoya U.)

in collaboration with

K. Takahashi, H. Kumamoto, N. Yonemaru

(Kumamoto U.)

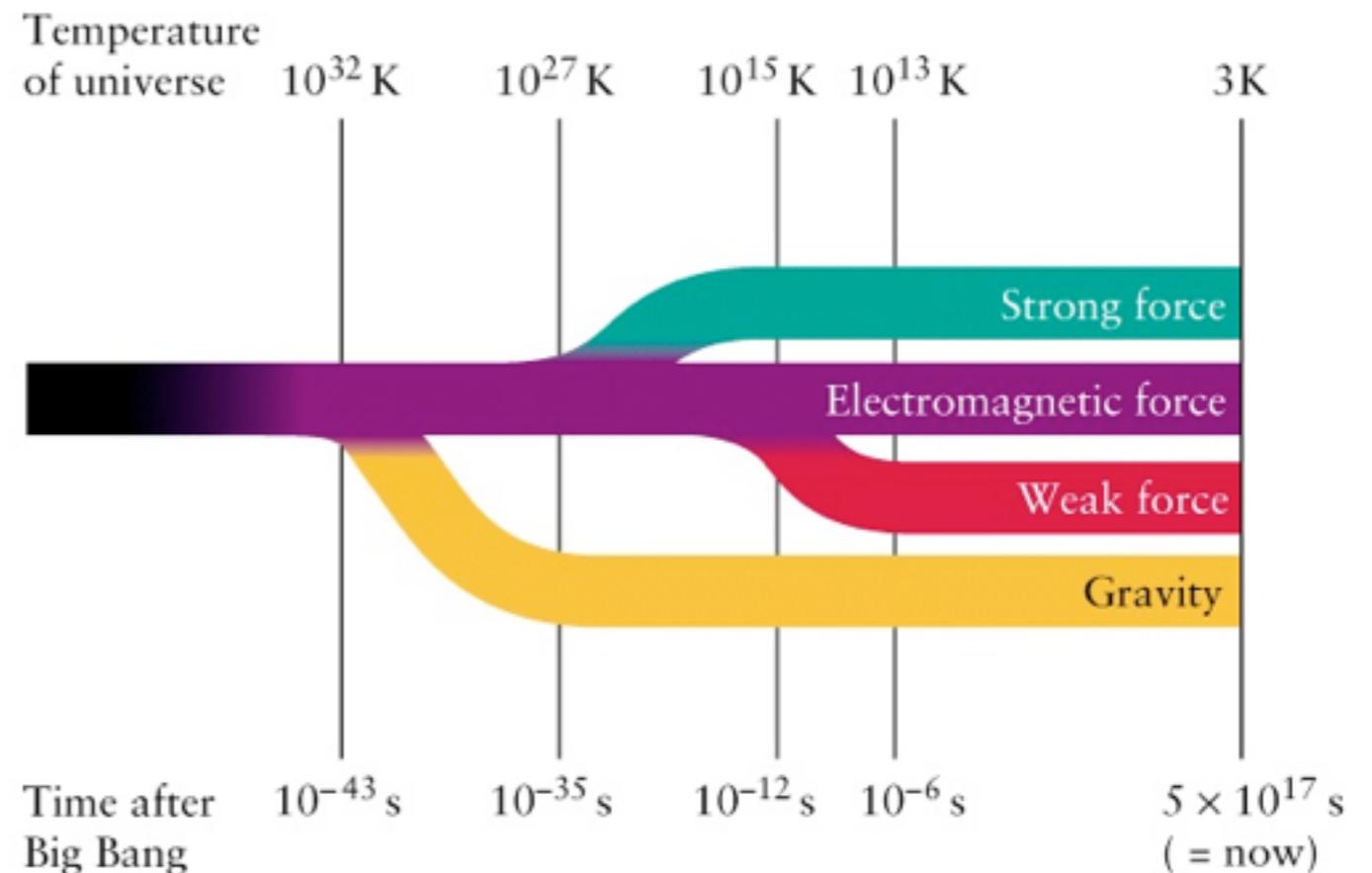
2016/4/8@KASI

Cosmic string ?

One dimensional topological defect generated in the early universe

Generation mechanism

1: Phase transition



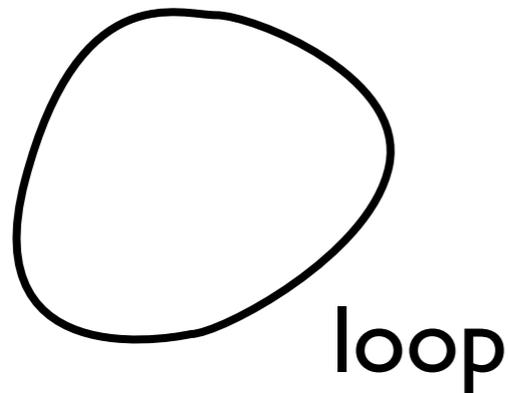
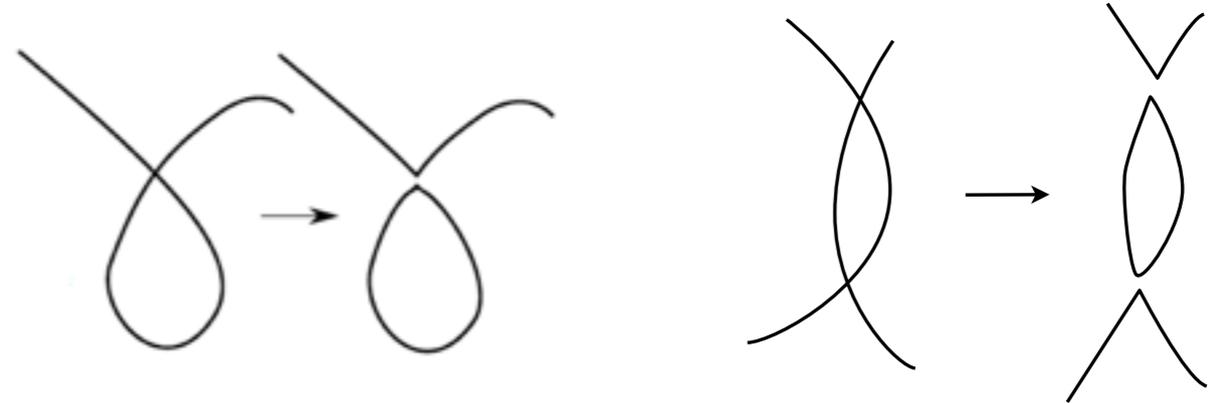
2: Cosmic superstrings

Cosmological size strings remains after inflation

→ could provide some insights into fundamental physics

Cosmic string ?

infinite string becomes a loop by **reconnection**

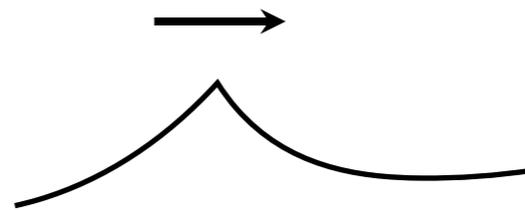


loop

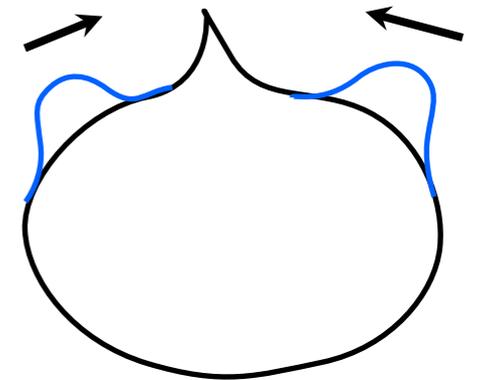
strings emit **gravitational waves**
especially from singular structures

infinite
string

kink

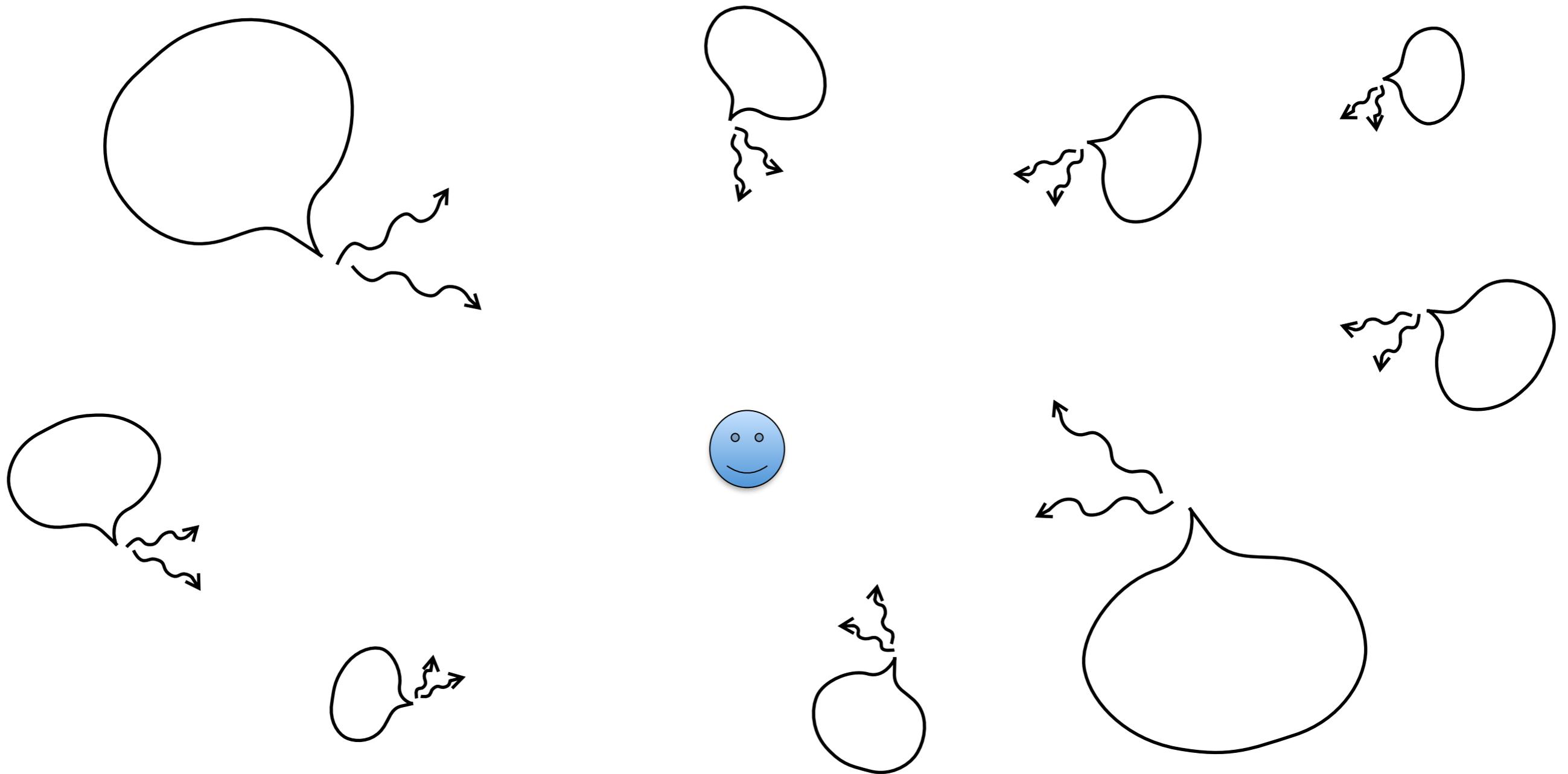


cusp



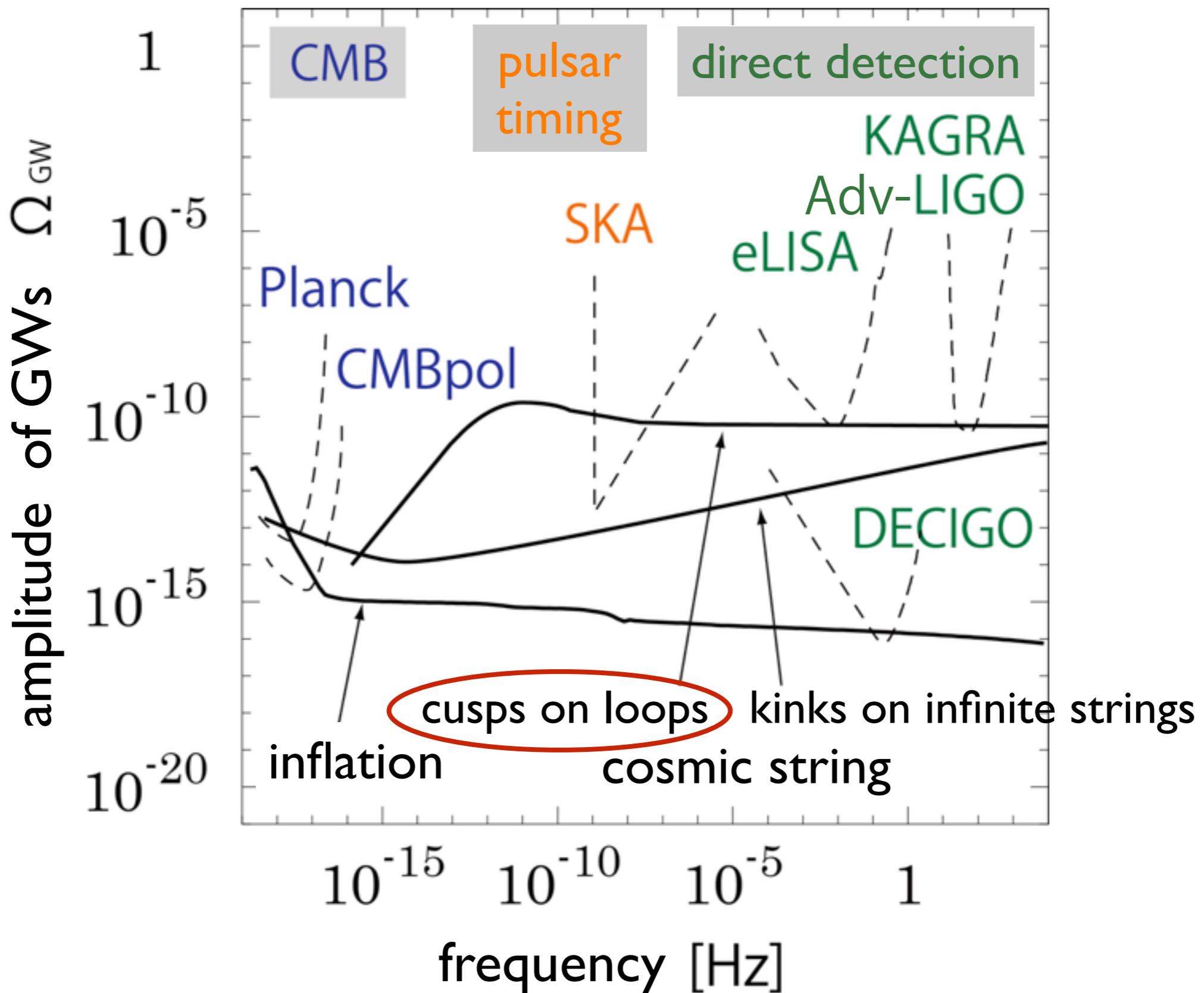
loops lose energy and shrink by emitting
gravitational waves and eventually **evaporate**

Gravitational waves from cosmic string loops



Gravitational waves coming from different directions overlap each other and form **gravitational wave background**

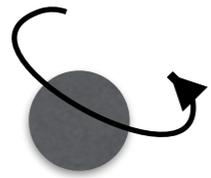
Spectrum and Sensitivities



Pulsar timing array

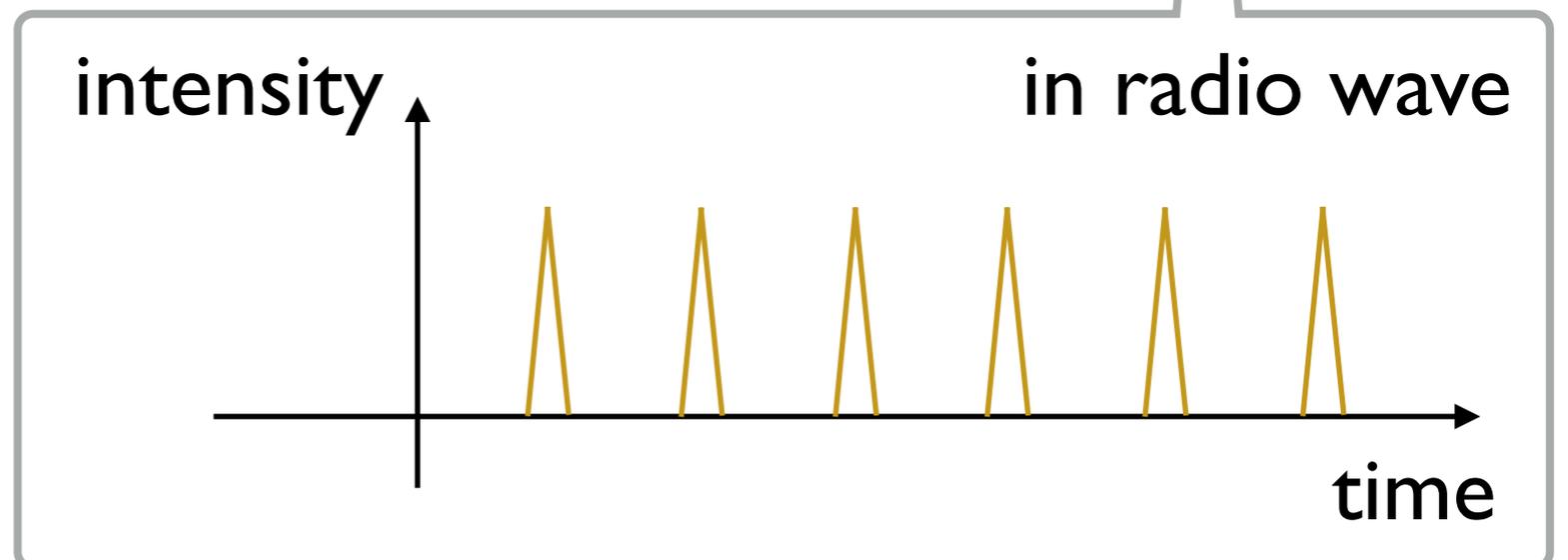
- Robust and unique test of gravitational waves

millisecond pulsar



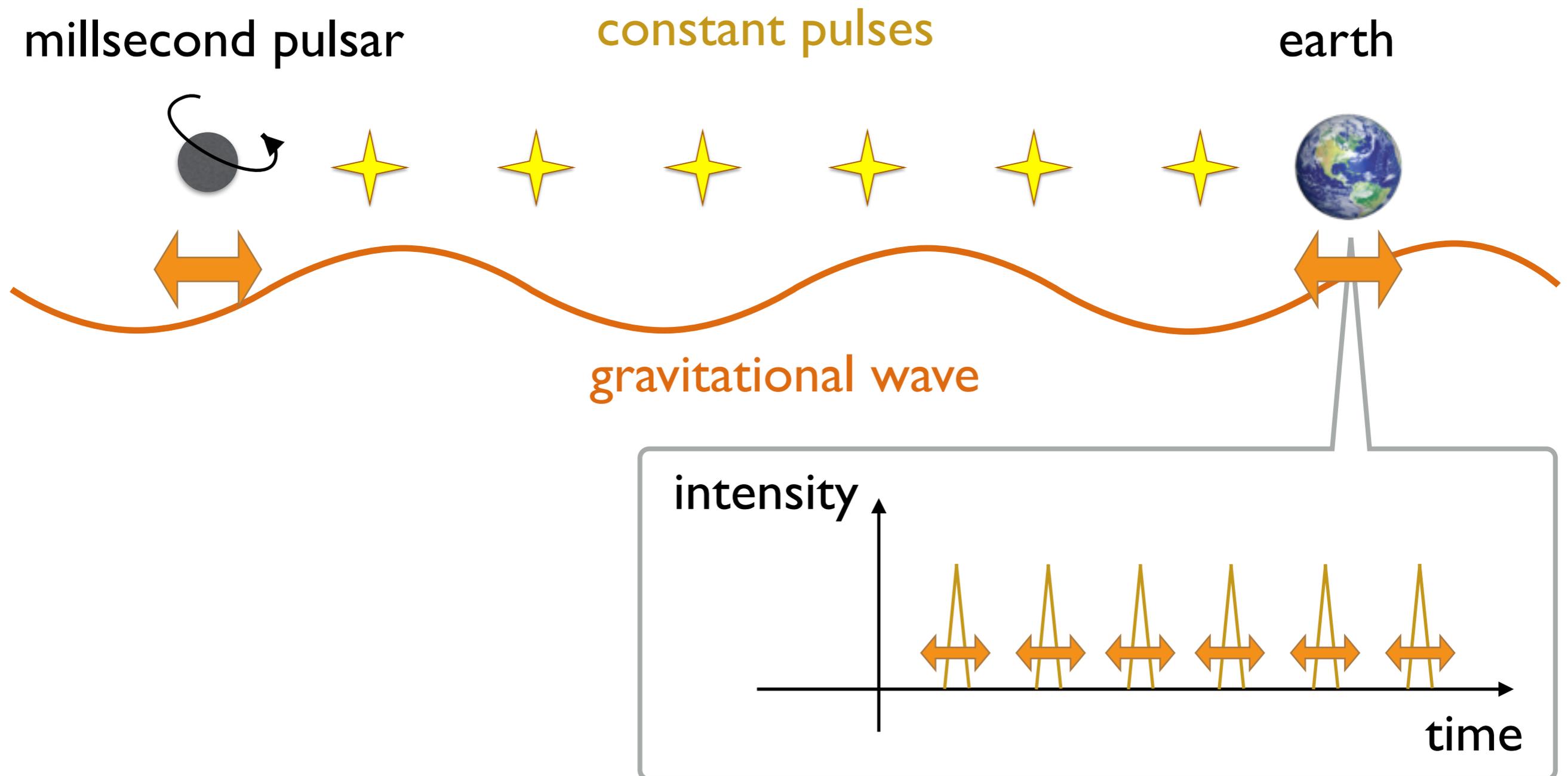
constant pulses

earth

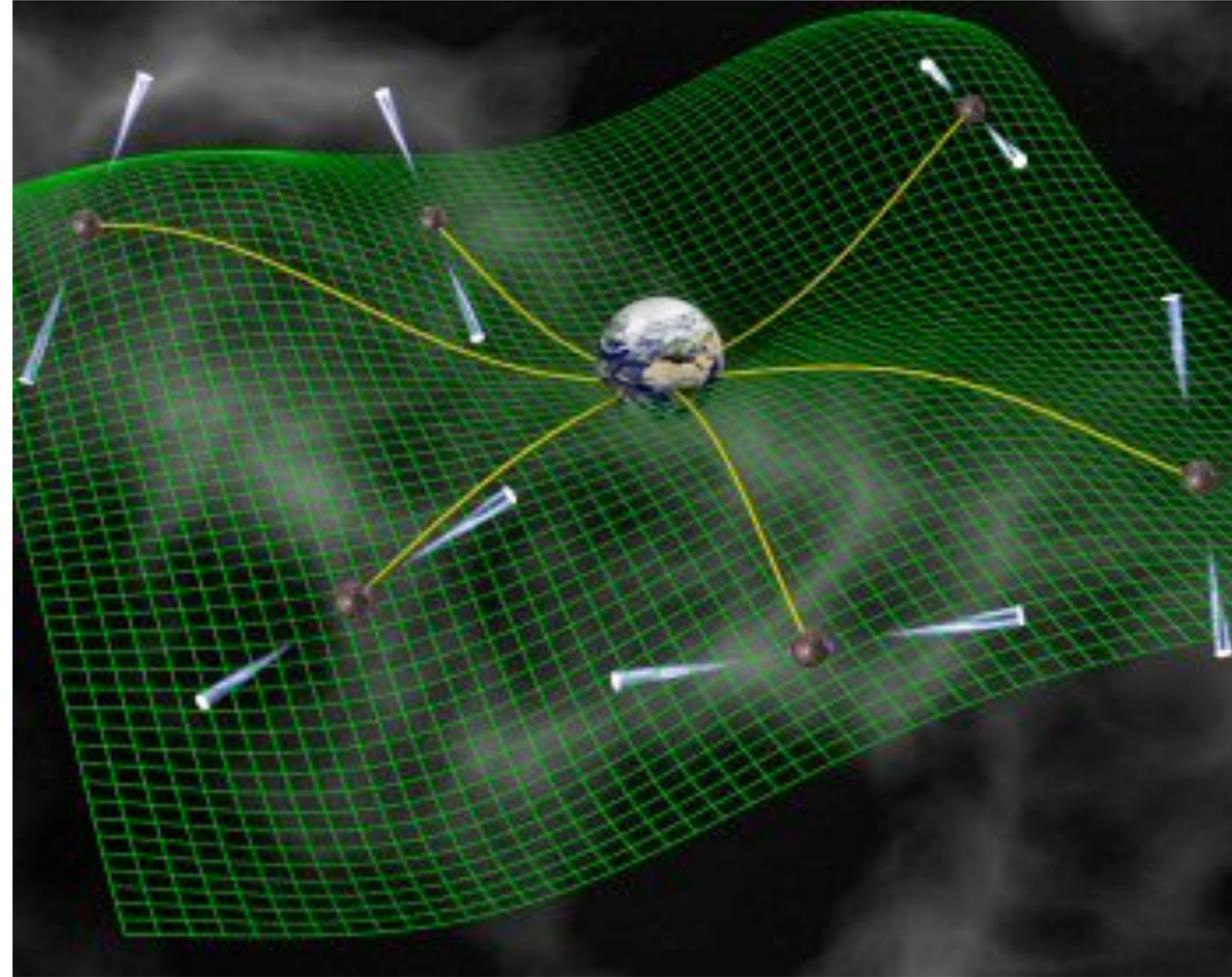


Pulsar timing array

- Robust and unique test of gravitational waves

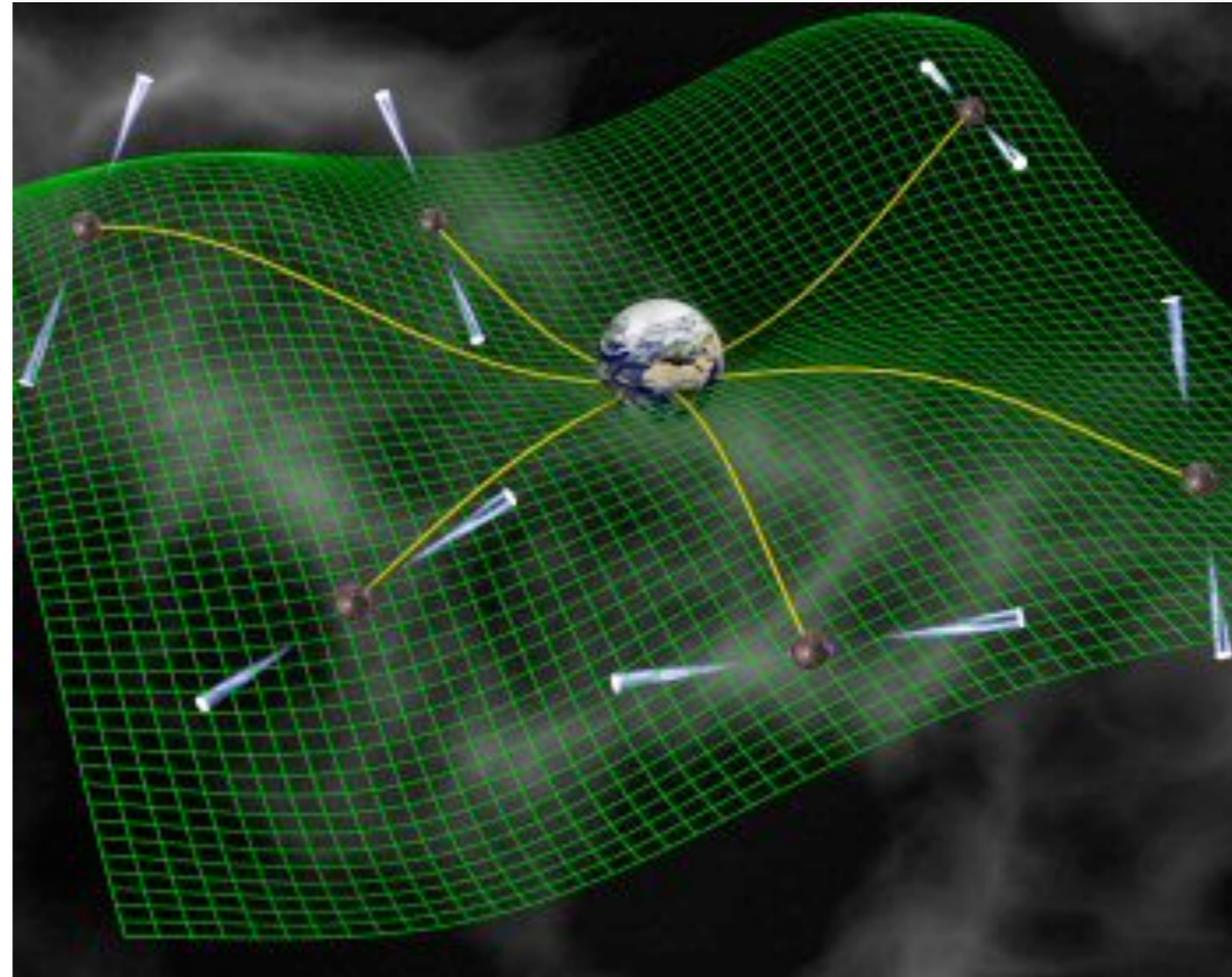


Pulsar timing array



- Correlation analysis with multiple pulsars to remove noises associating with individual pulsars
- Frequency \propto (Observation time)⁻¹

Pulsar timing array



- expected to detect gravitational waves from super massive black hole binaries or maybe from exotic sources (e.g. **cosmic strings**)

3 main parameters to characterize cosmic string

$G\mu$: tension = line density

Generation mechanism

determines

→ amplitude of a single burst

α : initial loop size $L \sim \alpha H^{-1}$

Network evolution

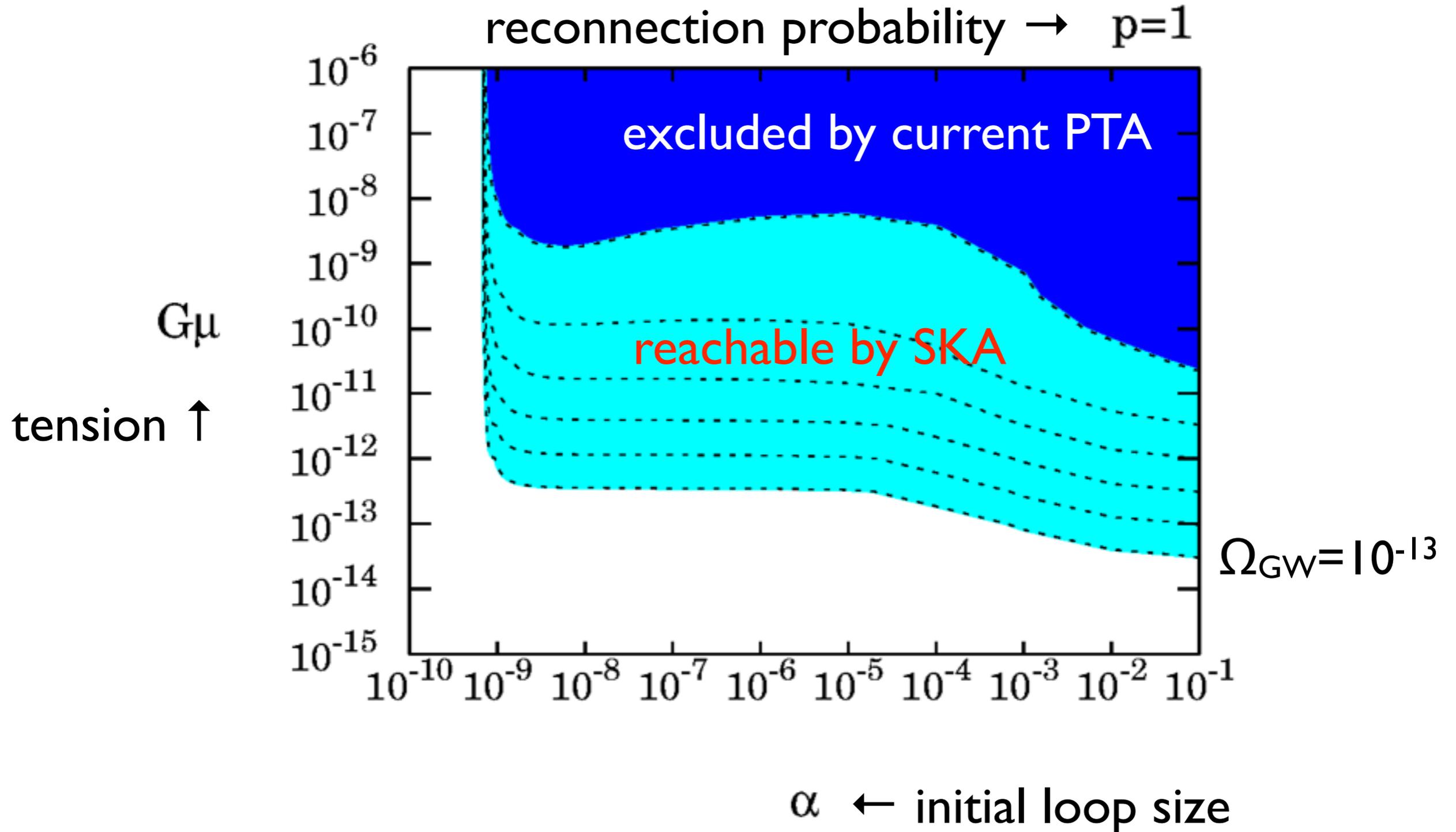
→ number density of loops

p : reconnection probability

Phase transition origin: $p=1$

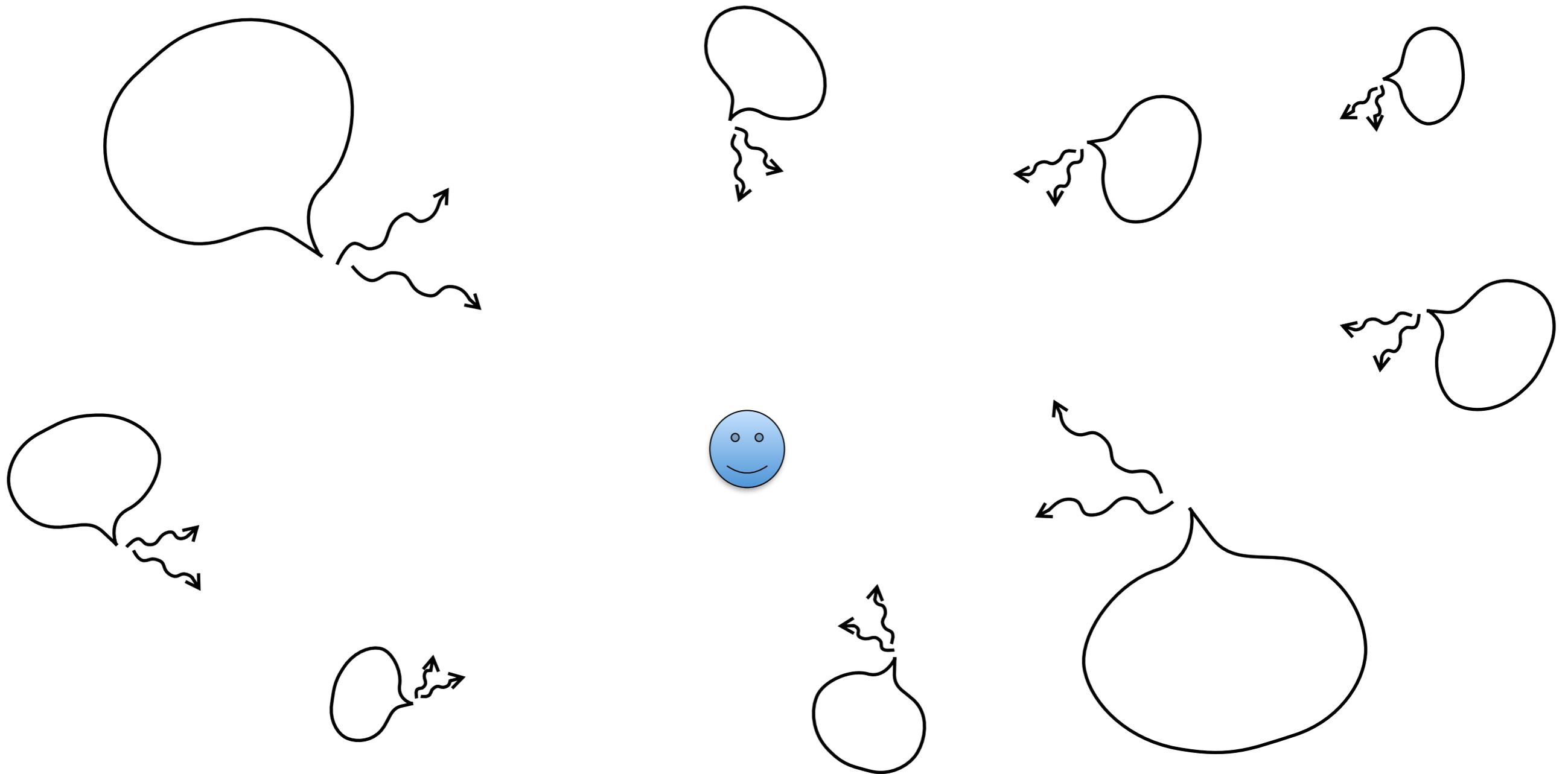
Cosmic superstring: $p \ll 1$

Detectability in $G\mu$ - α plane



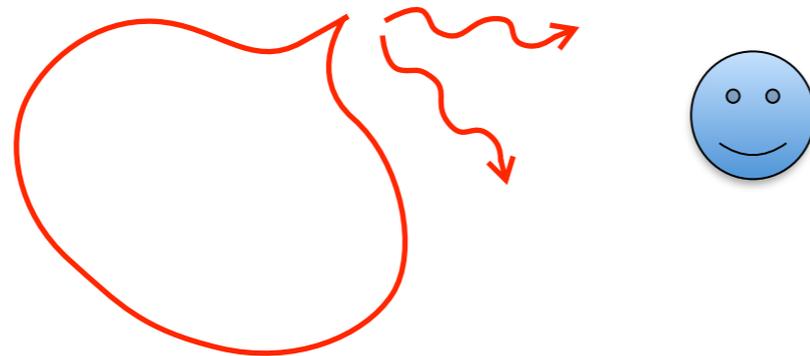
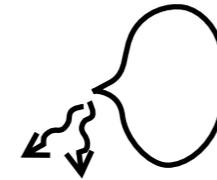
SKA will cover a large parameter space of cosmic string parameters

Gravitational waves from cosmic string loops



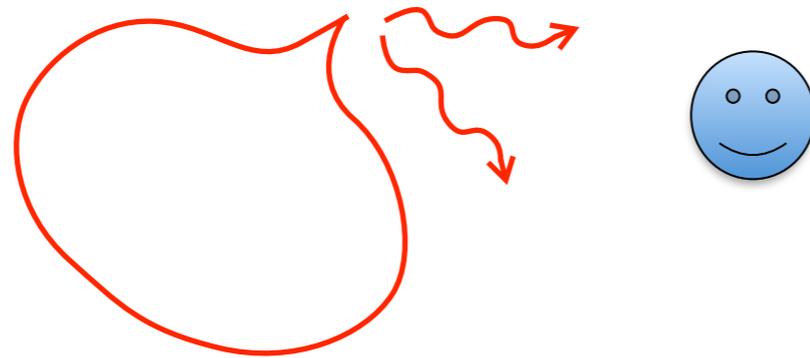
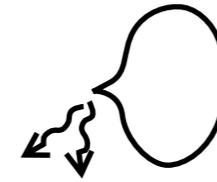
Gravitational waves coming from different directions overlap each other and form **gravitational wave background**

Anisotropy of gravitational wave background



If only a few loops contribute to the GW background,
it becomes **anisotropic**

Anisotropy of gravitational wave background



initial loop size α



We can extract information of loop number density from anisotropy of the gravitational wave background

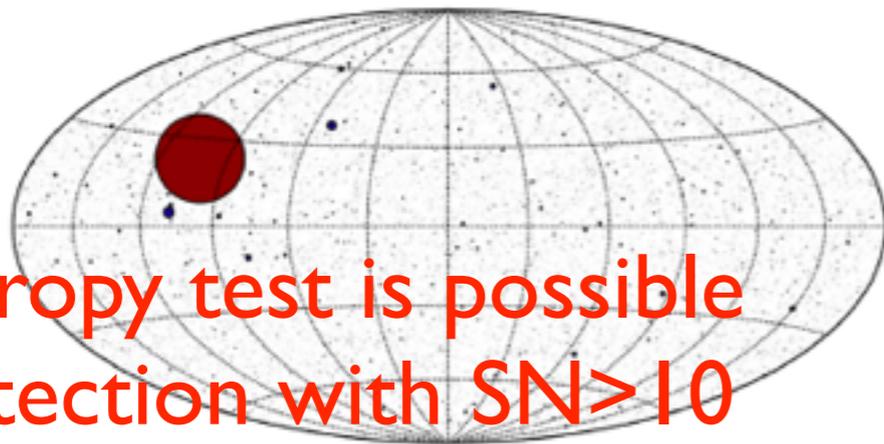
Anisotropy test

- Formulations are constructed by Mingarelli et. al., PRD 88, 062005 (2013)

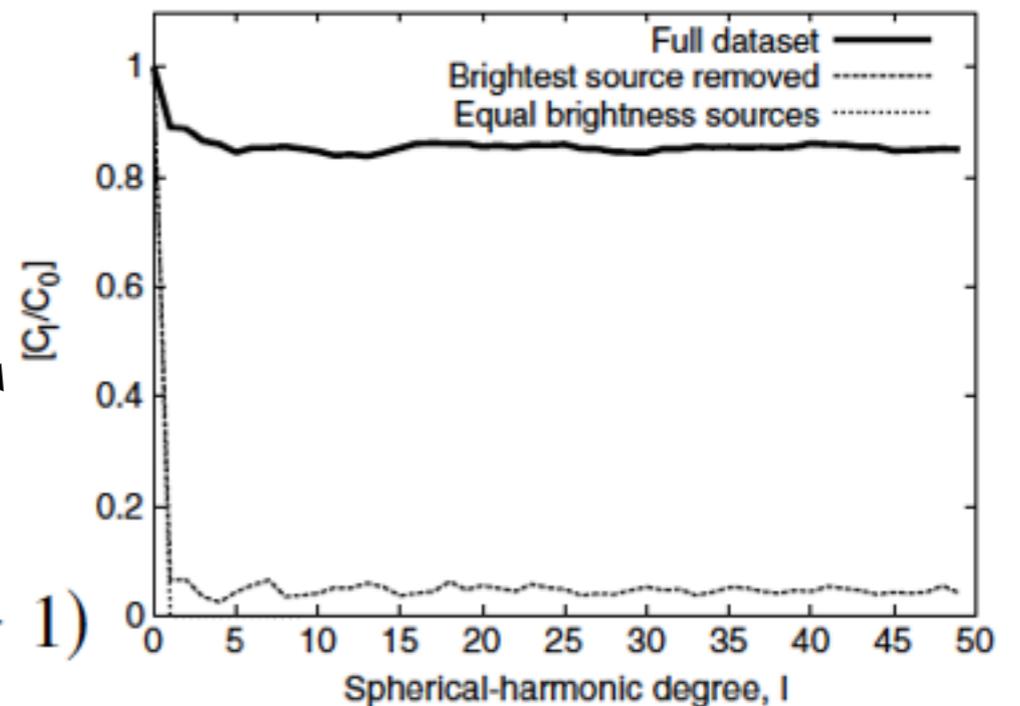
GW amplitude: $\Omega_{\text{gw}}(f) = \frac{8\pi^2}{3H_0^2} f^3 H(f) \int d\hat{\Omega} P(\hat{\Omega})$

$P(\hat{\Omega}) \equiv \sum_{lm} c_l^m Y_l^m(\hat{\Omega})$ ← Spherical harmonic expansion

- Simulation study in a context of GWs from SMBH binaries Taylor & Gair, PRD 88, 084011 (2013)



Anisotropy test is possible for detection with $\text{SN} > 10$

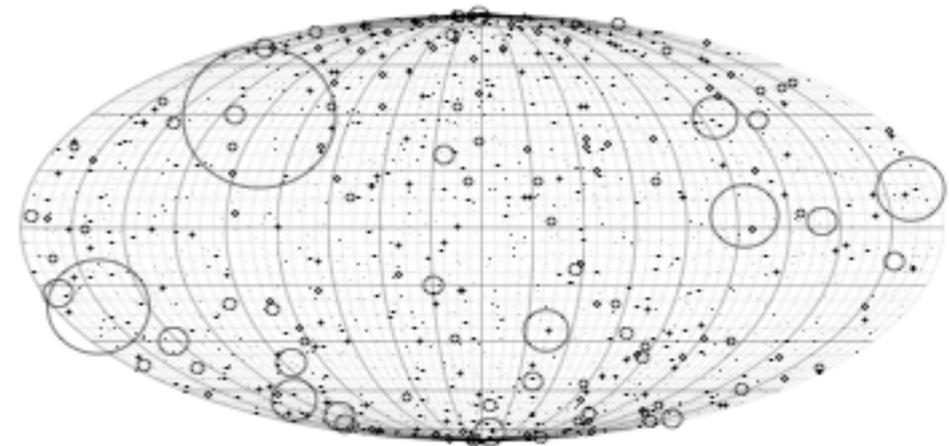
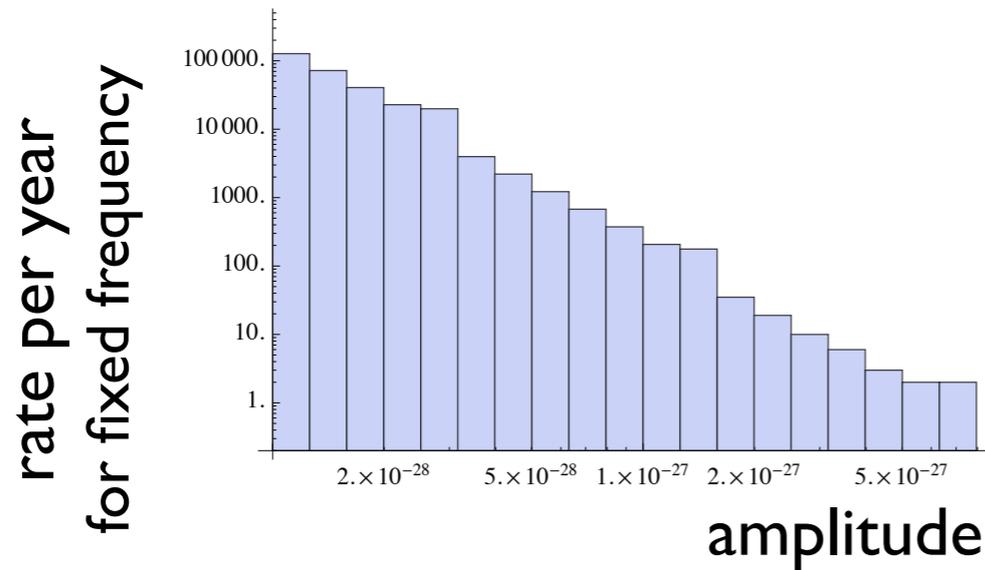


angular power spectrum: $C_l = \sum_m |c_{lm}|^2 / (2l + 1)$

Predictions for cosmic strings

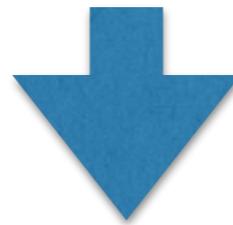
1. predict rate of GW bursts

2. randomly distribute the source



3. make 100 realizations

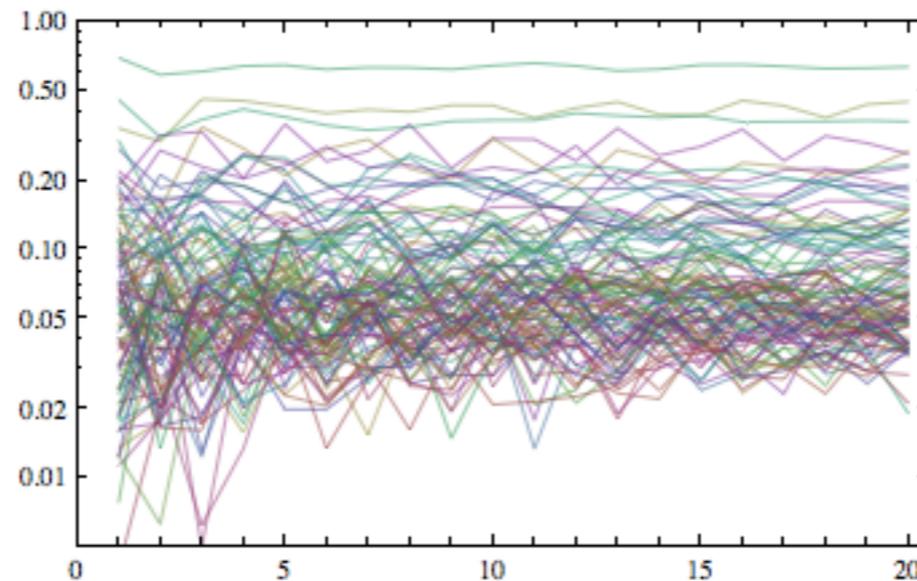
$$c_{lm} = \sum_{i=1}^N \rho_i Y_{lm}(\hat{\Omega}_i)$$



$$C_l = \sum_m |c_{lm}|^2 / (2l + 1)$$

4. calculate anisotropy levels

C_l / C_0

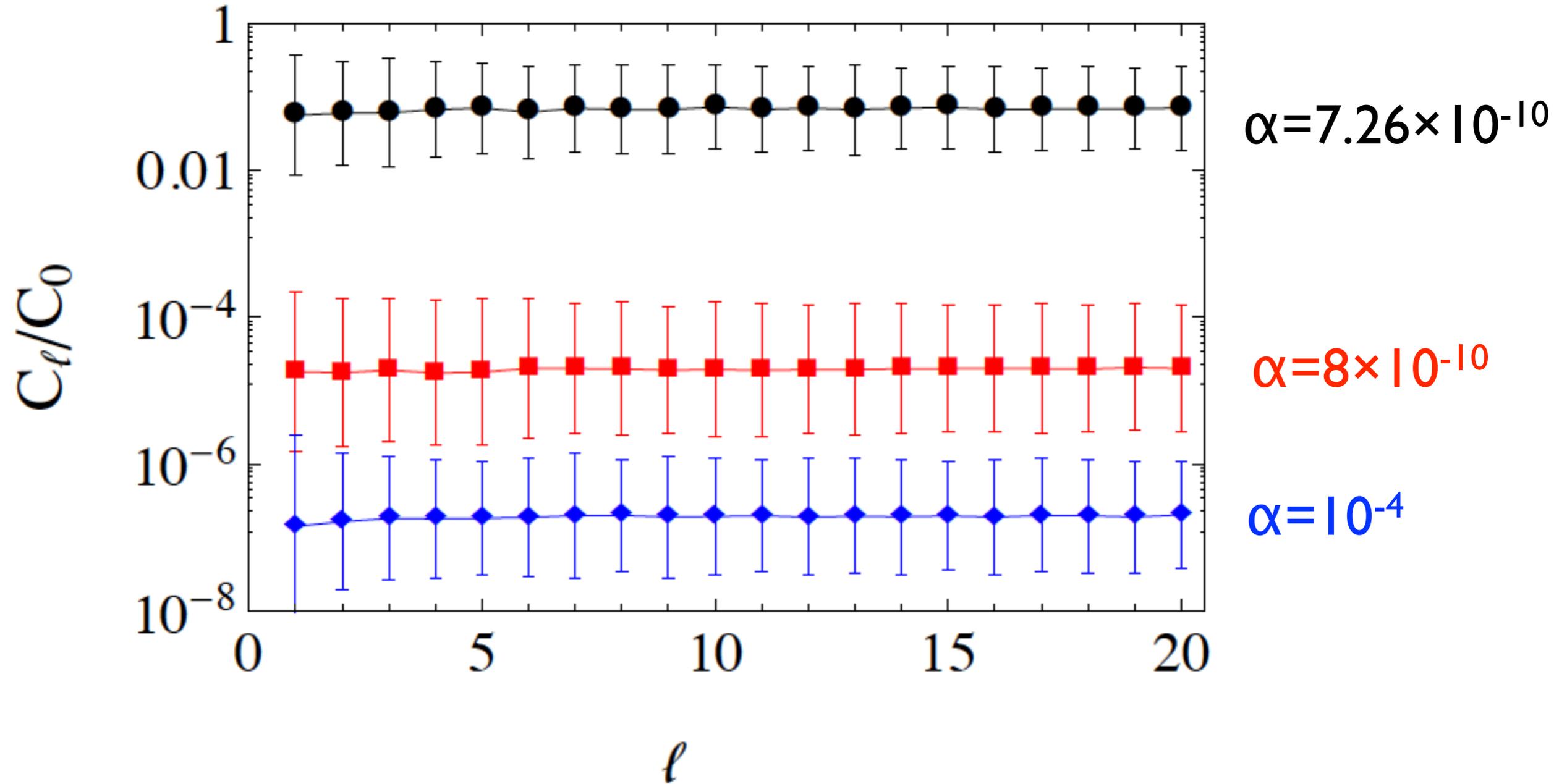


l : multipole

assumed $f=(10\text{yr})^{-1}$
 $G\mu=10^{-11}$, $p=1$

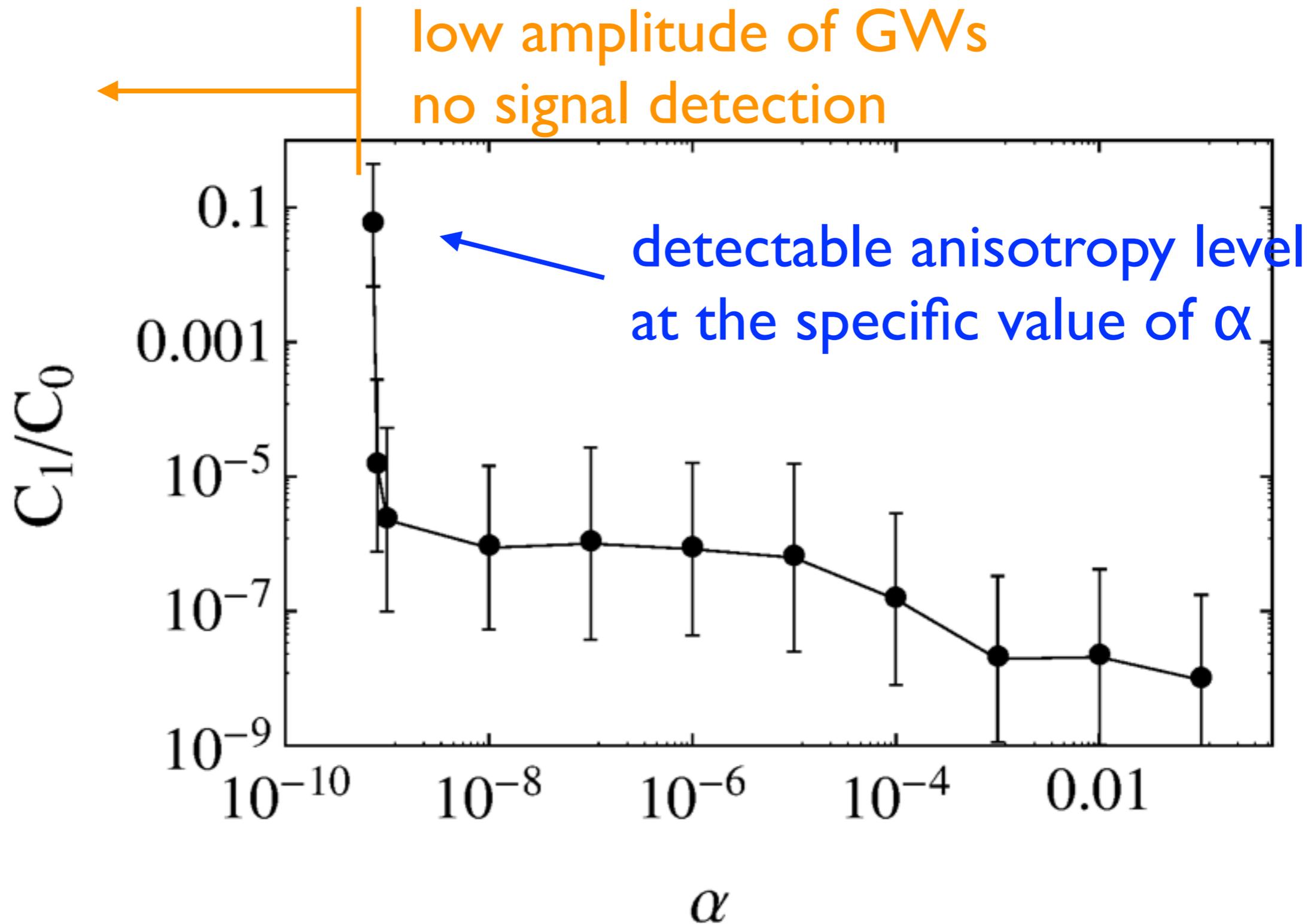
Anisotropy level

with 2σ error bar



We get large anisotropy for smaller value of α
(smaller initial loop size)

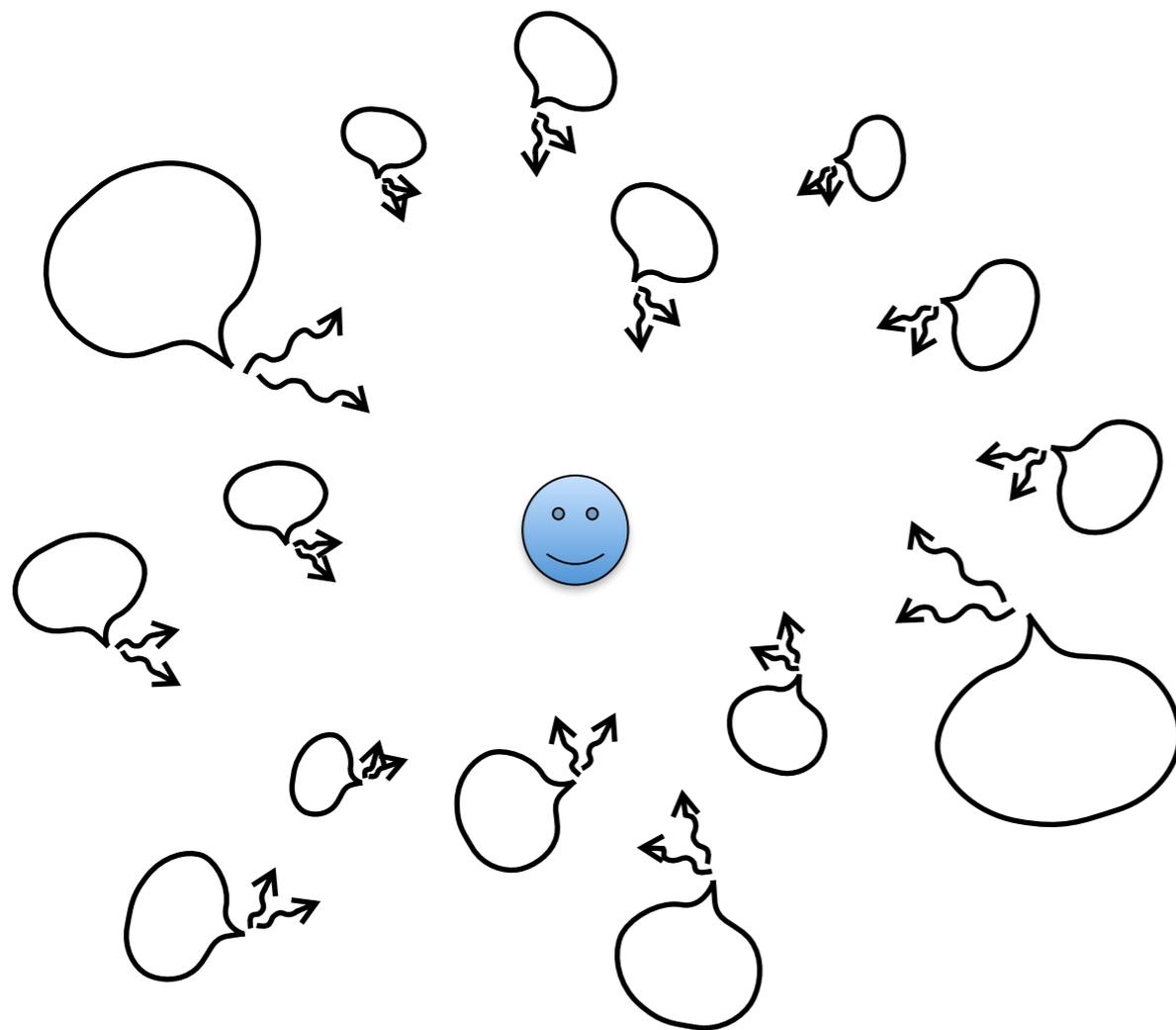
Dipole component α dependence



Interpretation

large α = long lifetime

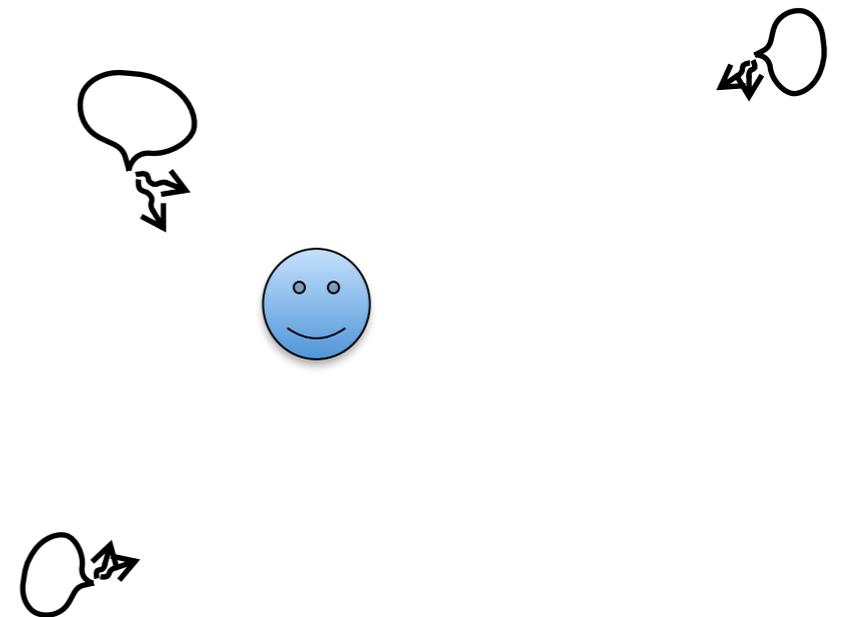
network consists of
mixture of old and new loops



→ isotropic GWs

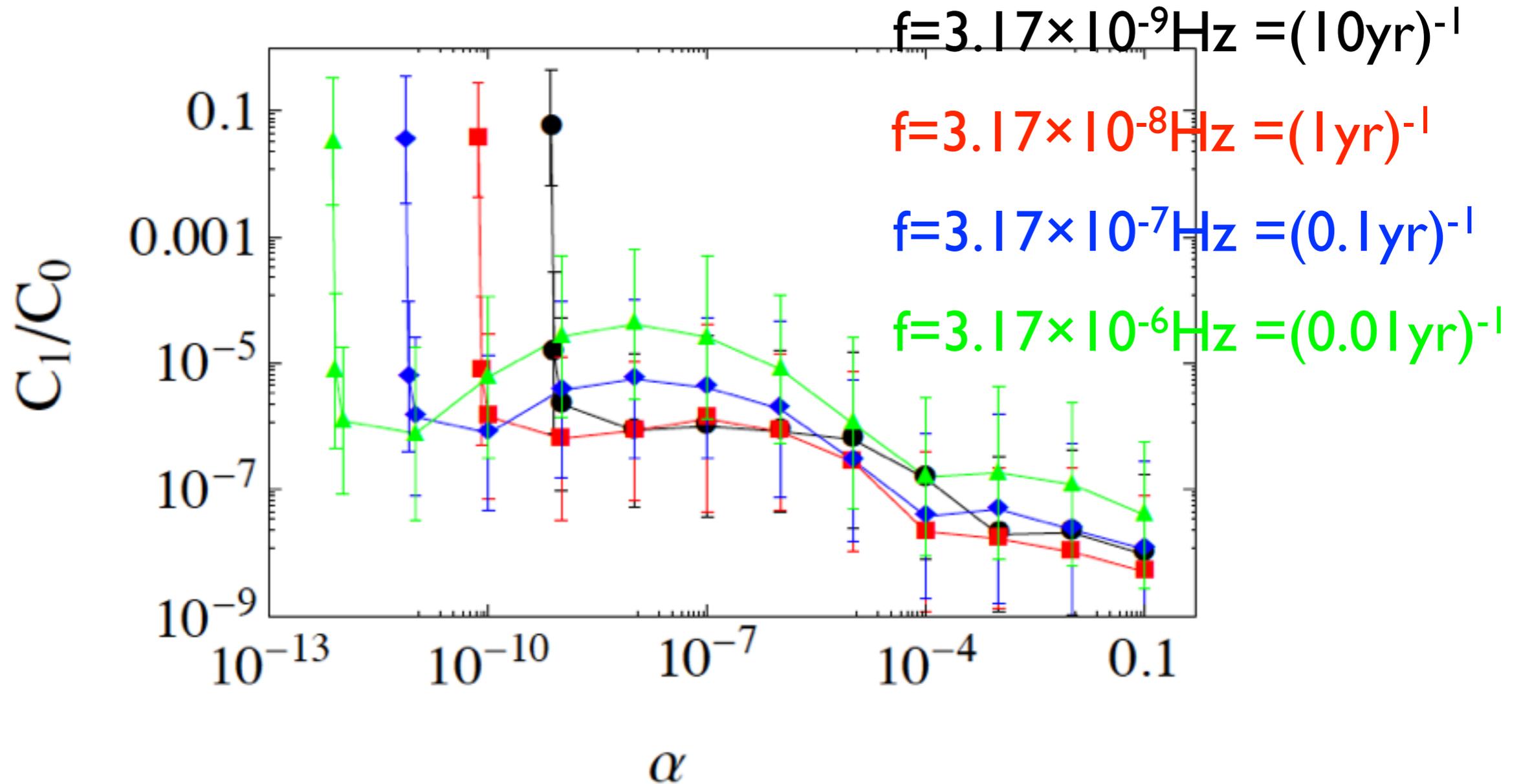
small α = short lifetime

loops evaporate soon after
their formation



→ anisotropic GWs

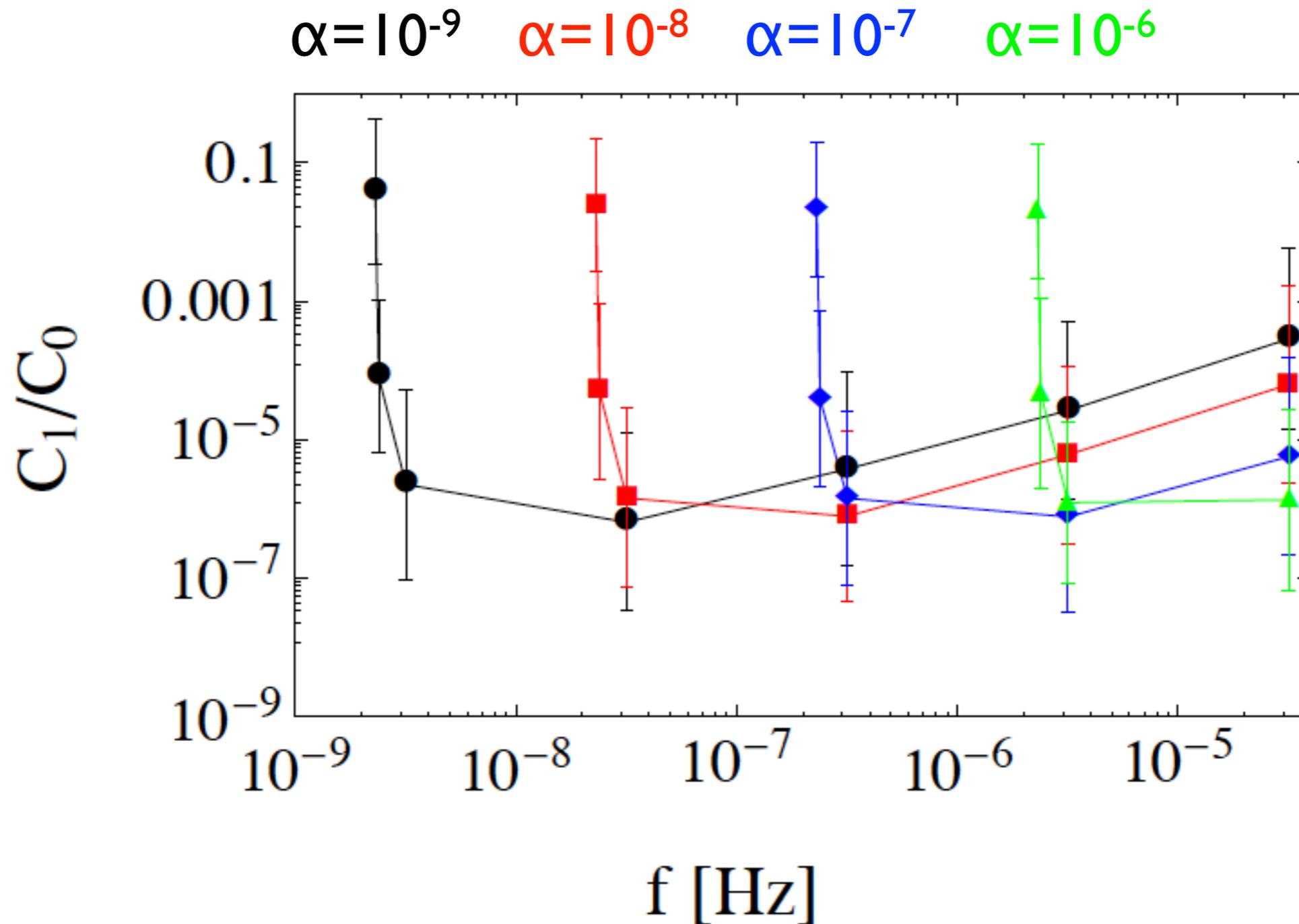
Observation frequency dependence



The peak position changes for different observation frequency

$$\text{frequency of GWs} \sim (\text{initial loop size})^{-1} \propto \alpha^{-1}$$

Observation frequency dependence



By checking anisotropy for different frequency bands, it may be possible to obtain implication on the value of α

Summary

- Testing the existence of cosmic string by PTA is important for obtaining implication on fundamental physics.
- SKA will cover a large parameter space of cosmic string parameters.
- Anisotropy of the gravitational wave background can be used to extract information on the initial loop size, which is important for understanding cosmic evolution of string network.